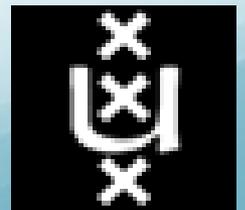


# The subquantum arrow of time

Theo M. Nieuwenhuizen  
University of Amsterdam



Emergent Quantum Mechanics 2013,  
EmQM13, Vienna



To understand Nature  
we have become accustomed  
to inconceivable concepts ...

*Our task is to demystify physics*

# Setup

New insights from quantum measurement theory  
The quantum measurement problem

Towards emergent QM  
On the (quantum) vacuum

Stochastic electrodynamics (SED)  
The subquantum arrow of time  
Bell inequalities

The H ground state in SED

# Fresh insights in Good Old QM

Allahverdyan, Balian, Nh, Physics Reports 2013:

“Understanding quantum measurement from dynamical models”

Solution of the Curie-Weiss model for Q-measurements:

*Unitary dynamics of tested system  $S$  + apparatus  $A$*

Results:

*Truncation* of the density matrix (*decay of Schrodinger cat terms*)

Fast, physical process, due to coupling to  $A$ .

*Registration*: Pointer of  $A$  goes to a stable state, triggered by the measured value

Amplification of small quantum signal due to initial metastability of  $A$

QM itself describes statistics of measurement outcomes;

*no measurement postulates needed; no extensions of QM needed*

Q-measurements lead to *statistical interpretation* of QM,

*frequency interpretation* of probabilities

# The measurement problem

(a problem for theorists and philosophers)

How to describe the individual events observed in practice?  
(How to go from *wave theory* to *events*?)

Quantum oddity: A *mixed density matrix* can be decomposed *in any basis*.  
Why would measurement basis be preferred?

“Unsolvable” => many interpretations: Copenhagen, many worlds, mind-body  
or extensions: spontaneous collapse models

Resolution by ABN'13:

Near the end of the measurement, *dynamical effects* in the apparatus  
make most decompositions of the density matrix *unstable*.

*Only the decomposition on the measurement basis is dynamically stable.*

So this is the physical basis. Arbitrary subensembles can be decomposed  
on this basis => connection to ordinary probabilities, frequency interpretation

# Towards emergent QM

In Nature: separate measurements occur

**We lack a theory that describes individual measurements**

Look for “subquantum mechanics”, “hidden variables theory”

This task is **more fundamental** than the search for quantum gravity,  
(and could have unpleasant surprises for it)

# On the (quantum) vacuum

The **Casimir effect** is a *real effect*

*Boats in harbours “attract each other” because few waves fit in between them*

Suppose: **Quantum vacuum** = **real** physical vacuum

**Zero point fluctuations** due to real fields, which **induce q-behavior**

**Up to which energy is the vacuum filled?** If not up to the Planck energy, *quantum gravity is useless, string theory can only be an effective theory*

**Picture:** vacuum fields gets created *after* the beginning of the Universe.

Maximal filling energy **below Planck energy** => **non-quantum** behavior **at Planck scale**

Vacuum energy (and pressure) are **borrowed from gravitation**.

**Cosmological constant** protected by **energy conservation**; fine tuning needed. (N'11)

**Particles are solitons**, affected by vacuum fluctuations

=> **Stochastic soliton mechanics** underlies quantum mechanics

# Stochastic Electrodynamics, SED

Vacuum = stochastic EM fields, energy per mode  $\frac{1}{2}\hbar\omega$ , spectrum  $\frac{\hbar\omega^3}{2\pi^2c^3}$

Classical theory, explains many quantum properties (talk Cetto)

Empty vacuum + SED spectrum = Lorentz invariant physical vacuum  
(Minkowski space-time + SED spectrum = Minkowski space-time)

This must explain *all* quantum behavior of atoms and molecules

*Zero* adjustable parameters, “*infinitely*” *many* constraints

# Example: the H atom in SED

Electron in classical **Kepler** orbits

It **radiates** away energy, would fall onto nucleus

It **absorbs energy** from fluctuating vacuum **EM fields**  
=> goes to other Kepler orbit. Statistics should produce  $\psi_0$ .

*If* there is a stable state, there is input+output of energy:  
**energy throughput**, current of energy to maintain stable state.

But this is an *arrow of time*

# The subquantum arrow of time

If there is a classical-type picture of the hidden variables theory, then a throughput of energy imposes quantum stability

**This implies an arrow of time**

This arrow is **more fundamental** than the thermodynamic and cosmological ones

# Bell Inequalities ??

Bell inequalities involve **non-commuting** variables

Hence these are measured one-by-one (Clauser, Aspect)

Next, they are inserted in an inequality **meant for commuting variables**

When the inequality is **violated**, it **cannot apply** to this situation:

There is a **contextuality loophole**, which **cannot be closed**

(Related to the detectors and the vacuum)

*The only conclusion is that QM works.*

*Not any implication on local realism.*

N'11

# SED: the H ground state

**Pro:** radiation and stochastic terms have desired scaling with  $\alpha$  and  $Z$   
*Leading logarithm of Lamb shift comes out in 2 lines*

**Contra:** *Fokker-Planck approximation* (2<sup>nd</sup> order in stochastic field) *fails*  
The theory is considered false, even by most advocates

Cetto & de la Pena: **resonances** appear beyond 2<sup>nd</sup> order, **induce q-behavior**

N'13: **Higher order corrections** in stochastic field, smaller by powers of  $\alpha$ ,  
develop **arbitrary powers of  $t$** , due to (higher order) resonances

**Conclusion:** *perturbation theory fails, the case is still open*

# Cole-Zou 2003: simulation of H ground state

Long box with  $L_x = L_y = 70a_0$ ,  $L_z \sim 10^6 a_0$

$\Rightarrow$  stochastic EM fields in *lowest x-y mode; many z-modes*

Periodic boundary conditions  $\Rightarrow$  linear spectrum,  $\omega = c|k_z|$

In atomic units  $\ddot{\mathbf{r}} = -\frac{\mathbf{r}}{r^3} + \mathbf{E} + \dot{\mathbf{r}} \times \mathbf{B} + \tau_e \ddot{\mathbf{r}}$   $\tau_e = \frac{2}{3}\alpha^3 \tau_0$

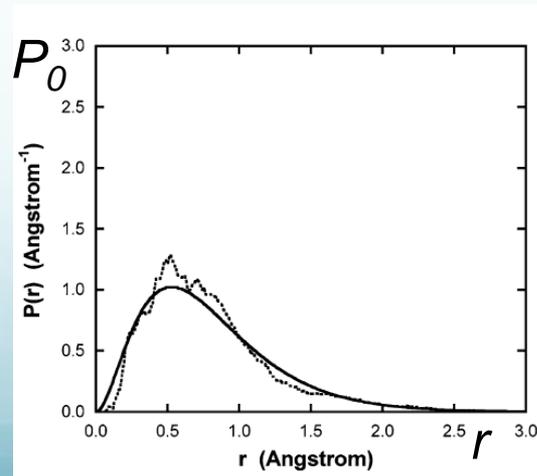
Neglect magnetic fields  $\Rightarrow$  motion in x-y plane

Resonances occur, they bring e to other Kepler orbits

Cole & Zou, 2003

*Encouraging similarity to quantum result*

$$P_0(r) = 4r^2 e^{-2r}$$



# Simulations anno 2013 at the University of A'dam



*M. Liska,  
E. van Heusden*

Solve in-plane motion up to  $10^5$  Bohr times

Remains cumbersome. Electron often evaporates or falls into nucleus

No definite conclusion reached

# But wait,

Coupling of  $e$  to EM fields shifts them; this generates the damping term.

*The damping is geometry dependent*

In long box:  
standard damping

$$\tau_e \ddot{\mathbf{r}} \rightarrow -\gamma \dot{\mathbf{r}}$$

$$\gamma = \frac{2\pi\alpha^3 Z^2}{\tau_0} \frac{a_0^2}{L_x^2} = \frac{9 \cdot 10^{-6}}{\tau_0}$$

situation  
Cole-Zou

Orbit remains in  $z=0$  plane  $\ddot{\mathbf{r}} = -\frac{\mathbf{r}}{r^3} + \mathbf{E} - \gamma \dot{\mathbf{r}}$

Numerically: problems remain

# Protocol for H ground state

Consider the nearly-conserved quantities

*E = energy*

*L = angular momentum of in-plane motion*

*$\lambda$  = angle of Runge-Lenz vector*

Integrate them analytically over one orbit, iterate this numerically

*Work in progress*

# What are we looking for?

*2d H ground state*  $\psi_0^2(\mathbf{r}) = \frac{8}{\pi} e^{-4r}$

In classical approach with **weak noise**: *density in phase space* =  $f(E, L)$

$$P_0(\mathbf{p}, \mathbf{r}) = \frac{8}{\pi^2 E^2} e^{4/E} \quad \int d^3p P_0(\mathbf{p}, \mathbf{r}) = \frac{8}{\pi} e^{-4r} \quad N'05$$

Parameters Kepler orbits  
distributed uniformly in  $L$

$$P(E, L) = \frac{32\sqrt{2}}{|E|^{7/2}} e^{4/E}$$

# Summary

*QM does not describe individual measurements  
They do occur, so an underlying less-statistical theory must exist*

*Local, classical picture may underlie quantum mechanics  
Many constraints, no free parameters*

*Bell inequalities do not rule that out,  
contextuality loophole cannot be closed*

*Atomic stability then implies a “subquantum” arrow of time;  
more fundamental than thermodynamic and cosmological arrows*

*Structure of H ground state in SED is studied, work in progress*

# Quantum mechanics of hydrogen atom: nucleus charge = $-Ze$

Spectrum: Rydberg energy

$$\alpha^2 Z^2 mc^2$$

Relativistic corrections

$$\alpha^4 Z^4 mc^2$$

Lamb shift

$$\alpha^5 Z^4 mc^2 \log \alpha Z$$

Relativistic spectrum for  $m=c=1$

$$E_{n,l} = \left\{ 1 + \frac{\alpha^2 Z^2}{\left[ n - l - \frac{1}{2} + \sqrt{\left( l + \frac{1}{2} \right)^2 - \alpha^2 Z^2} \right]^2} \right\}^{-1/2} \approx 1 - \frac{\alpha^2 Z^2}{2n^2} - \alpha^4 Z^4 \left[ \frac{1}{n^3(2l+1)} - \frac{3}{8n^4} \right]$$

Lamb shift: not from Schrodinger equation,  
but due to coupling to EM field

weak effect  $\Rightarrow$  weak coupling, weak Lorentz damping

# Weak damping classical stochastic theories for hydrogen atom

Phase space density

$$P(\mathbf{r}, \mathbf{p}, t) = \langle \delta(\mathbf{r}(t) - \mathbf{r}) \delta(\mathbf{p}(t) - \mathbf{p}) \rangle$$

Evolution

$$\partial_t P = -\mathcal{L}P + \text{Lorentz damping} + \text{diffusion terms}$$

$$\mathcal{L} \equiv \frac{1}{\gamma} \mathbf{p} \cdot \nabla_{\mathbf{r}} - \frac{\mathbf{r}}{r^3} \cdot \nabla_{\mathbf{p}}$$

Stationary distribution = function of conserved quantities

Energy

$$H = \sqrt{1 + \alpha^2 Z^2 \mathbf{p}^2} - \frac{\alpha^2 Z^2}{r} = \gamma - \frac{\alpha^2 Z^2}{r}$$

Angular momentum

$$\mathbf{L} = \mathbf{r} \times \mathbf{p}$$

# The unsquared dance

Define  $R(E)$  by

$$E = \sqrt{1 + \frac{\alpha^4 Z^4}{R^2} - \frac{\alpha^2 Z^2}{R}} = \left(1 + \frac{2\alpha^2 Z^2}{RE}\right)^{-1/2}$$

Then non-relativistic problem

$$\frac{1}{2}\mathbf{p}^2 - \frac{\alpha^2 Z^2}{2r^2} - \frac{E}{r} = -\frac{E}{R}$$

Effective angular momentum

$$\omega = \sqrt{1 - \frac{Z^2 \alpha^2}{L^2}}, \quad \omega^2 L^2 = L^2 - \alpha^2 Z^2$$

In QM: effective angular momentum

$$\ell_1(\ell_1 + 1) = l(l + 1) - \alpha^2 Z^2, \quad \ell_1 = -\frac{1}{2} + \frac{1}{2} \sqrt{(2l + 1)^2 - 4\alpha^2 Z^2} \approx l - \frac{\alpha^2 Z^2}{2l + 1}$$

# Bits and pieces

Go to cylindrical coords

$$p \Rightarrow (R, \mu, \nu)$$

$$\omega L = r \sqrt{2E \left( \frac{1}{r} - \frac{1}{R} \right)} \sin \mu, \quad p_r = \sqrt{2E \left( \frac{1}{r} - \frac{1}{R} \right)} \cos \mu$$

Volume element in p-space

$$dV_p = \frac{d\mu d\nu dR}{2R^2 \Phi(E)} \sqrt{\frac{2}{Er} - \frac{2}{ER}} \sin \mu \quad \Phi(E) = \frac{\sqrt{1 + \alpha^4 Z^4 / R^2}}{2E^2}$$

Consider

$$\mathcal{P}(E, L) = \omega L R^3 \Phi(E) \exp(-\alpha R)$$

Then

$$dV_p \mathcal{P} = d\mu d\nu dR (R - r) e^{-\alpha R} \sin^2 \mu$$

Momentum integral

$$\int_{\mathbf{p}} dV_p \mathcal{P} = \frac{\pi^2}{\alpha^2} e^{-\alpha r}$$

= non-relativistic  
groundstate density

and the ratio

$$\frac{\omega^2 L^2 R}{2E} = r(R - r) \sin^2 \mu$$

Generates a factor  $r$

# Yrast states: $l=n-1$ (maximal angular momentum)

Phase space densities

$$\mathcal{P}(R, L) = C_{\omega} L R^3 \Phi(E) \left( \frac{\omega^2 L^2 R}{2E} \right)^{2\ell_1} e^{-2R E_{n_1}/(\ell_1+1)}$$

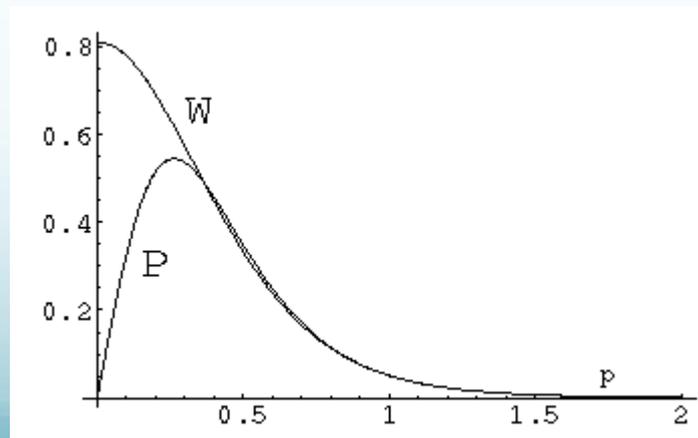
Momentum average gives square of wavefunctions:

$$\int_p dV_p \mathcal{P} = \int_{R,\mu,\nu} dV_p \mathcal{P} = \frac{2^{2+2\ell_1} E_{n_1}^{3+2\ell_1}}{(1+\ell_1)^{4+2\ell_1} \Gamma(2+2\ell_1)} r^{2\ell_1} e^{-2r E_{n_1}/(1+\ell_1)}$$

$n=1$ : Ground state:  $\mathcal{P}$  positive, so  $\mathcal{P}$  differs from Wigner function

Reason: our  $p$  is instantaneous; in Wigner function it is statistical

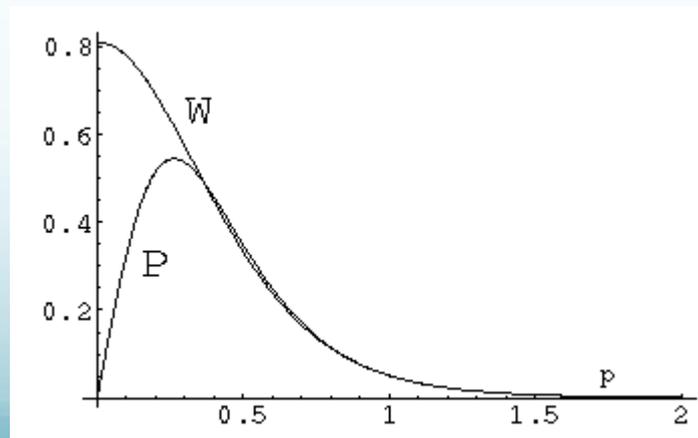
Space average



# Wigner(p) versus Phase space density(p)

Test: scatter fast electrons on hydrogen atoms  
(Mott & Massey: *Impulse approximation*)

Slow speeds: many revolutions during scattering: quantum cloud  
Fast speed: instantaneous position and speed of bound e is probed



$> \alpha c$

## Average energy

$$\langle E \rangle_{nl; cl} = 1 - \alpha^2 Z^2 \left\langle \frac{1}{R} \right\rangle + \frac{1}{2} \alpha^4 Z^4 \left\langle \frac{1}{R^2} \right\rangle = 1 - \frac{\alpha^2 Z^2}{2} - \frac{(4n^2 + 4n - 1) \alpha^4 Z^4}{4n^4 (8n^2 - 6n + 1)} = E_{nl} + \frac{\alpha^4 Z^4}{8n^4 (4n - 1)}$$

not correct ....

## Doing the forbidden: Neglect correlations

Approximate  $\left\langle \frac{1}{R^2} \right\rangle \Rightarrow \left\langle \frac{1}{R} \right\rangle^2$  then quantum mechanical energy recovered at order  $\alpha^4$

Do this at all orders  $\left\langle \frac{1}{R^k} \right\rangle \Rightarrow \left\langle \frac{1}{R} \right\rangle^k$

$$\langle\langle E \rangle\rangle_{nl; cl} \equiv \sqrt{1 + \alpha^4 Z^4 \left\langle \frac{1}{R} \right\rangle^2} - \alpha^2 Z^2 \left\langle \frac{1}{R} \right\rangle = \sqrt{1 + \frac{\alpha^4 Z^4 E_{nl}^2}{4(1 + \ell_l)^4}} - \frac{\alpha^2 Z^2 E_{nl}}{2(1 + \ell_l)^2}$$

Exact quantum result regained for Yrast states:

$$\langle\langle E \rangle\rangle_{nl; cl} = \sqrt{\frac{(1 + E_{nl}^2)^2}{4E_{nl}^2}} - \frac{1 - E_{nl}^2}{2E_{nl}} = E_{nl}$$

## 2p state: spherical harmonics

In frame along  $r$ , cylindrical coordinates:  $L$  involves angles  $\mu$  and  $\nu$

$$\mathbf{r} = r(\sin \theta \cos \phi, \sin \theta \sin \phi, \cos \theta), \quad \mathbf{L} = \rho r \sin \mu \hat{\mathbf{L}},$$

$$\hat{\mathbf{L}} = (-\cos \theta \cos \phi \sin \nu - \sin \phi \cos \nu, -\cos \theta \sin \phi \sin \nu + \cos \phi \cos \nu, \sin \theta \sin \nu)$$

Search phase space forms  $Y_{|m_2|}^* Y_{|m_2|} \rightarrow \mathcal{Y}_{|m_2|}^{|m_2|}$

Proposal:  $\mathcal{Y}_{1-1}^{1-1} = \frac{3}{4\pi}(\hat{L}_z^2 - \hat{L}_z), \quad \mathcal{Y}_{10}^{10} = \frac{3}{4\pi}(1 - 2\hat{L}_z^2), \quad \mathcal{Y}_{11}^{11} = \frac{3}{4\pi}(\hat{L}_z^2 + \hat{L}_z)$

a)  $\int_0^{2\pi} \frac{d\nu}{2\pi} \mathcal{Y}_{11}^{11} = \int_0^{2\pi} \frac{d\nu}{2\pi} \mathcal{Y}_{1-1}^{1-1} = \frac{3}{8\pi} \sin^2 \theta = |Y_{1\pm 1}|^2 \quad \int_0^{2\pi} \frac{d\nu}{2\pi} \mathcal{Y}_{10}^{10} = \frac{3}{4\pi} \cos^2 \theta = |Y_{10}|^2$

b)  $\overline{\hat{L}_z \mathcal{Y}_{1m;1m}} \equiv \int \sin \theta d\theta d\phi \int \frac{d\nu}{2\pi} \hat{L}_z \mathcal{Y}_{1m}^{1m} = m = \langle Y_{1m} | L_z \text{ op} | Y_{1m} \rangle, \quad (m = -1, 0, 1)$

## *Discussion*

Considered class of theories includes Stochastic Electrodynamics

Phase space densities proposed for Yrast states  $l=n-1$

Integral over  $p$  gives QM density

Integral over  $r$  does NOT give result from Wigner function

Test by scattering fast electrons on H

Different method, same result: consistency

Also 2s state considered: works in the same approach (non-unique)

$l=1$  phase space forms for squares of spherical harmonics proposed

Ground state density positive; excited states partially negative

Quantum energies recovered iff correlations neglected

Physically: time scale separation :

each new quantum operator corresponds to a classical average

at a well separated time  $\Leftrightarrow$  subensembles de la Pena & Cetto

## Theo's dream

- 1) Schrodinger mechanics = SED *de la Pena, Cetto, Cole, Khrennikov, ..*
- 2) Particles, photons: solitons in electro-gravity *Carter, Pereira, Arcos, Burinskii*
- 3) Physical explanation for exclusion principle and QM-statistics *timescales,*
- 4) QM = statistics of stochastic soliton mechanics *energetics*

*This dream integrates basically all works of Albert Einstein.*

*Now you may say I'm a dreamer  
But I'm not the only one  
I hope one day you'll join us  
And the world will be as one*

Imagine, John Lennon

# Quantum mechanics is a theory

that describes  
the statistics  
of outcomes  
of experiments

*It cannot and should not describe individual experiments  
(otherwise than in a probabilistic sense)*

